

Status of the purification at KamLAND

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CENBG CNRS/IN2P3 and Université Bordeaux I

LRT workshop, Aussois (France) October, 1-4 2006



KamLAND Collaboration

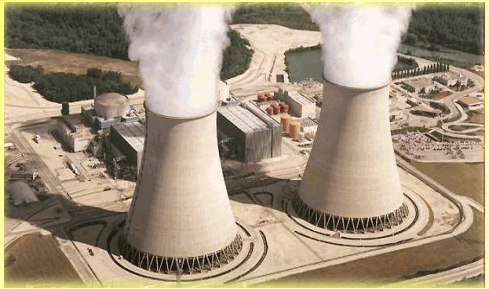


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California Institute of Technology, USA
University Bordeaux 1, France,
Drexel University, USA,
IHEP, China,
Kansas State University, USA,
Triangle Universities Nuclear Lab., USA,**

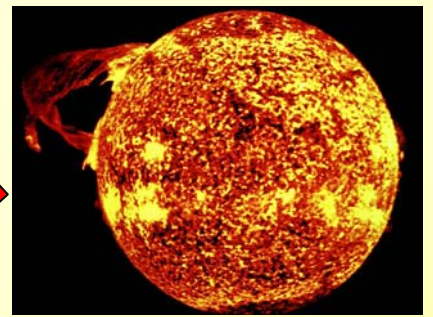
**University of Alabama, USA,
University of Hawaii, USA,
University of New Mexico, USA,
University of Tennessee, USA,
Lawrence Berkeley National Lab., USA,
Louisiana State University, USA,
Stanford University, USA**



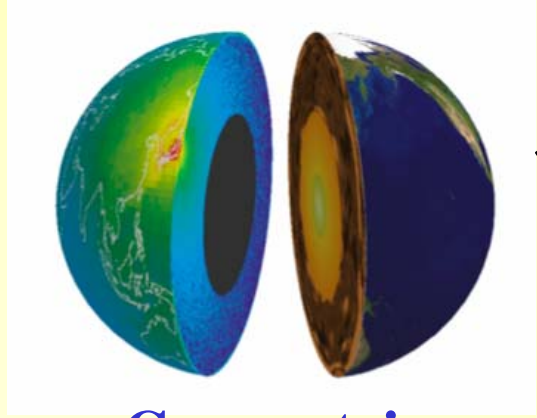
KamLAND Physics



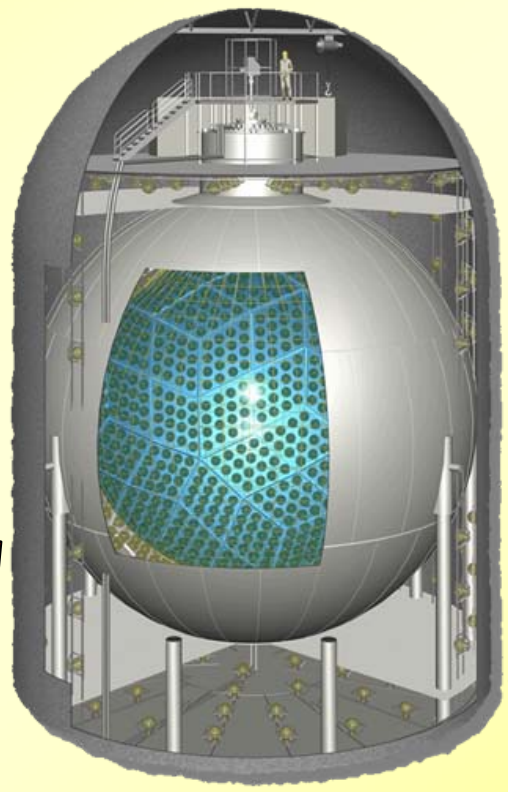
Reactors



Solar neutrinos



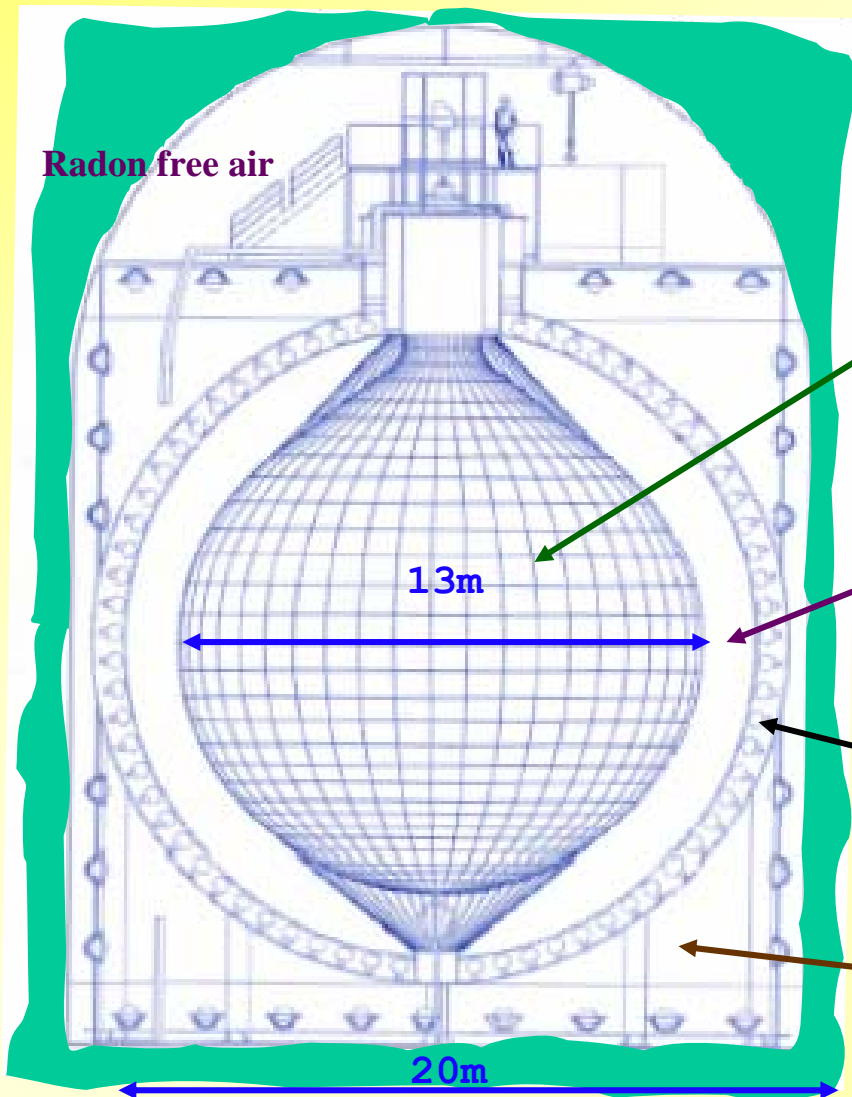
Geoneutrinos



Supernovae neutrinos



KamLAND Detector



**Kamioka mine overburden: 2700 w.m.e.
Muon rate: 0.34 Hz**

**1000 tons of liquid scintillator
80% dodecane+ 20% pseudocumene + PPO**

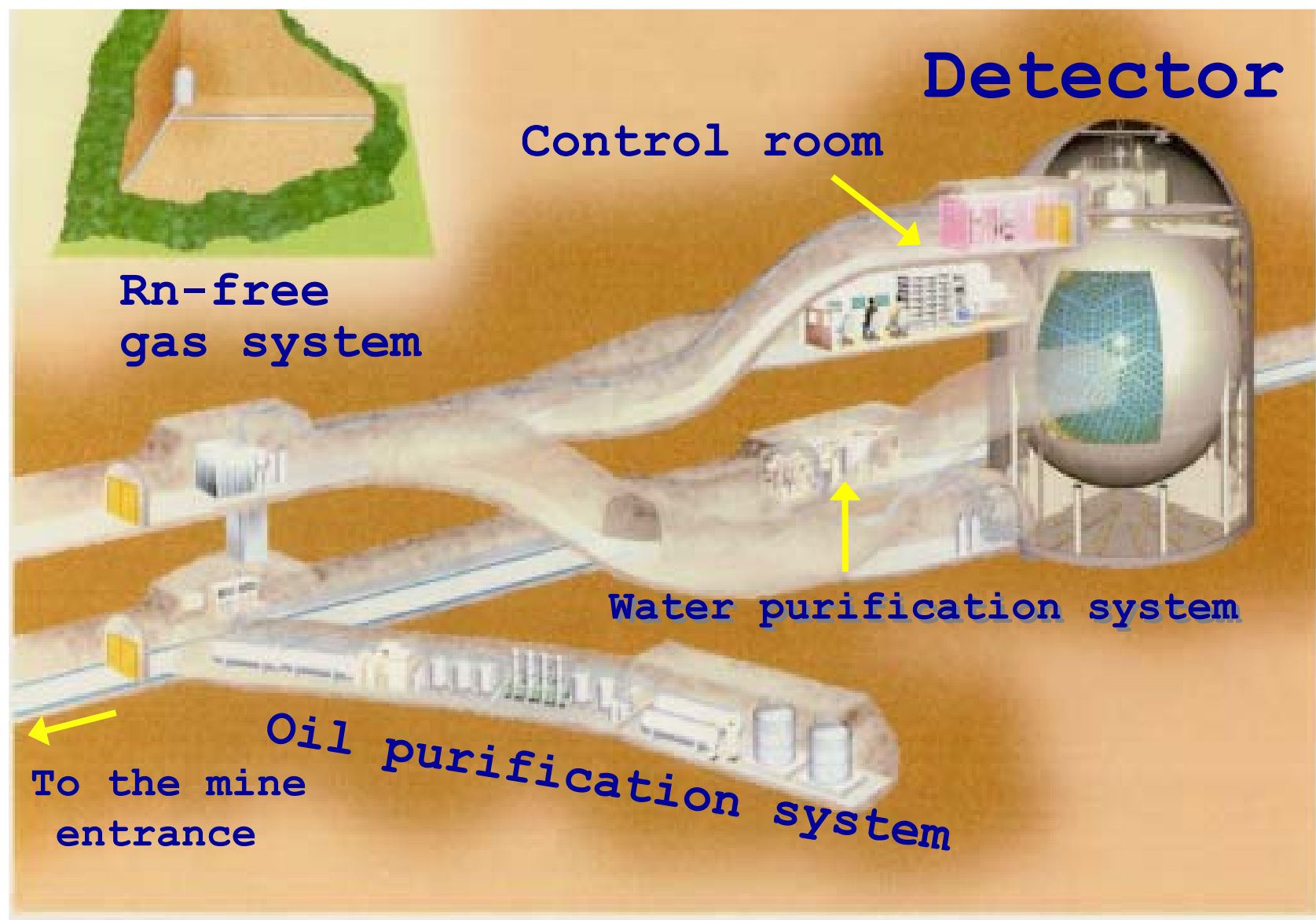
**Mineral oil
Buffer against external radiation**

**1979 PMTs (1325 17" + 554 20"
Photocathode coverage 34%**

**Outer water Cherenkov detector
for μ veto**



KamLAND installation



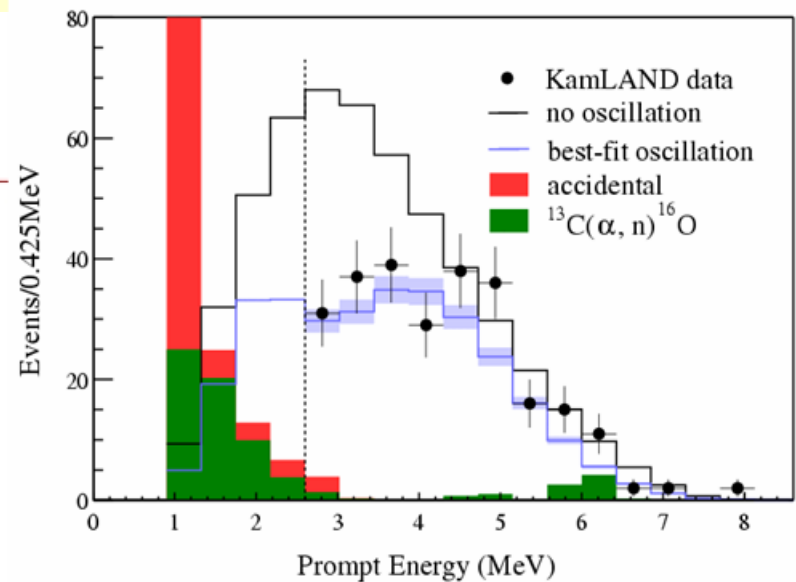
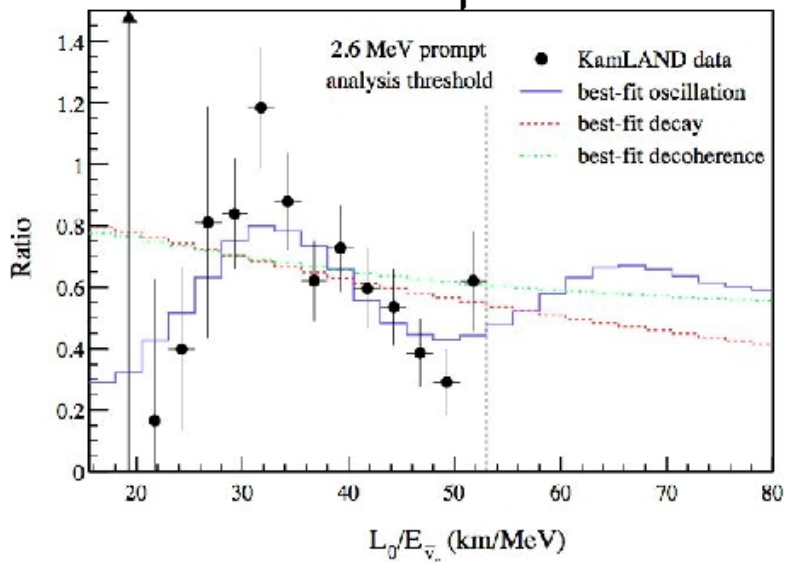


Reactor Anti-Neutrino Results

Phys.Rev.Lett. 94, 081801 (2005)

2nd result : disappearance
confirmed at 99.998% CL

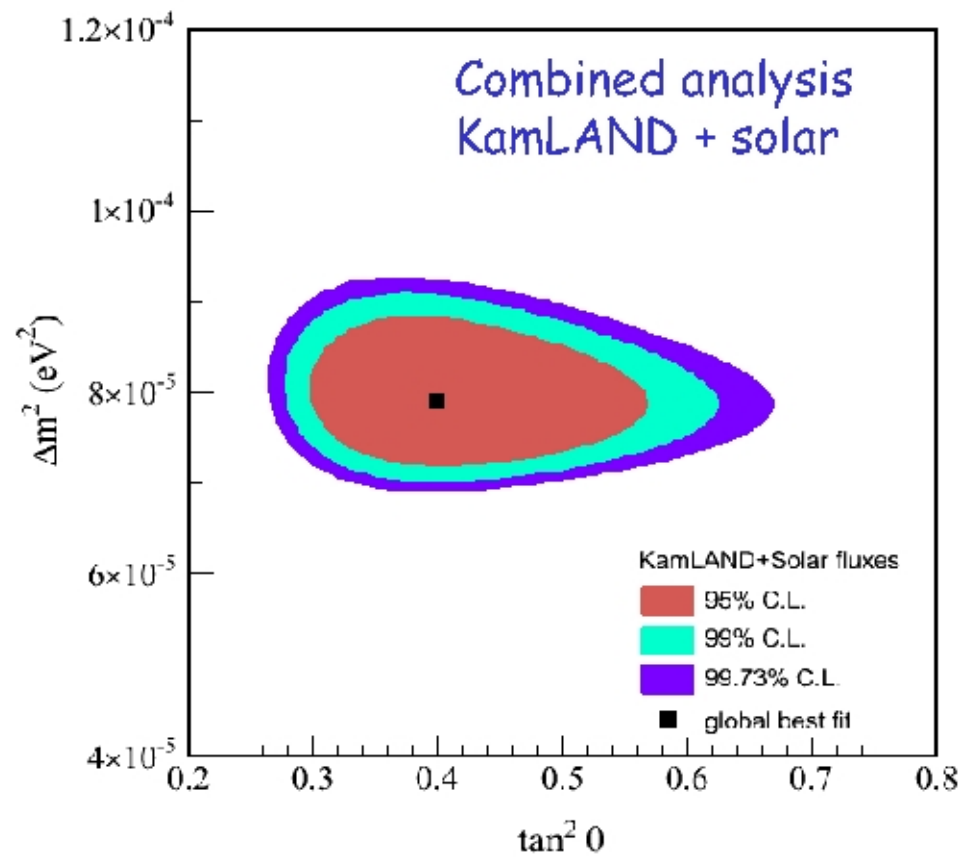
L/E plot



Scaled reduced spectrum
rejected @ 99.6 % CL

Eliminate other disappearance
hypothesis at 98% CL

Reactor Anti-Neutrino Results



Best fit :

$$\Delta m^2 = 7.9^{+0.6}_{-0.5} \times 10^{-5} \text{ eV}^2$$

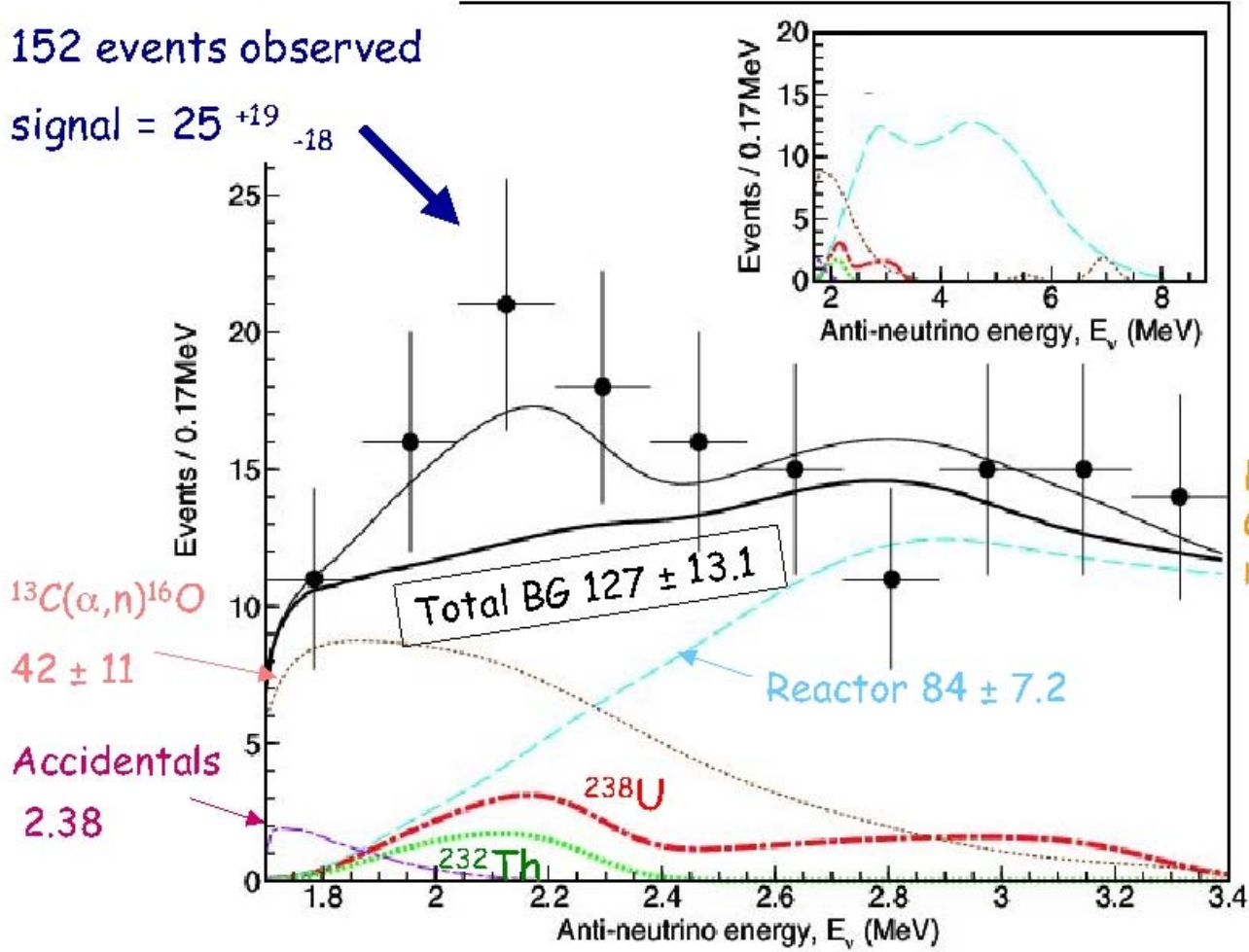
$$\tan^2 \theta = 0.40^{+0.10}_{-0.07}$$



Geoneutrinos Results

152 events observed

signal = 25^{+19}_{-18}



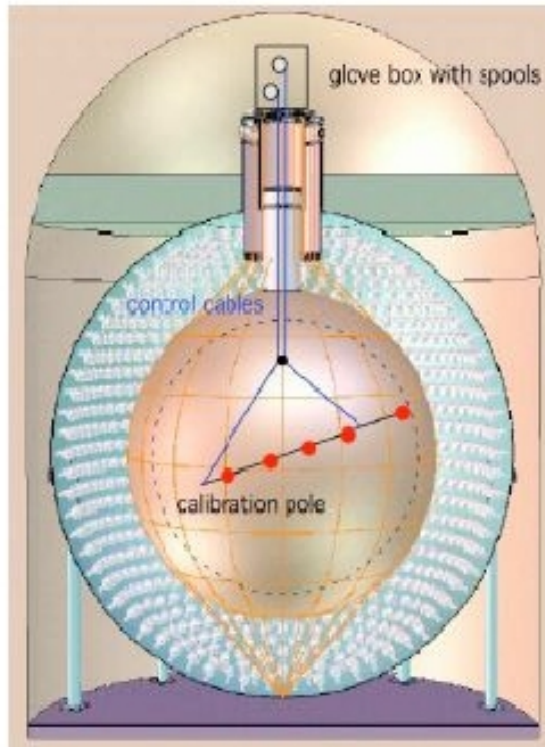
Data :
749.1 days
Mar 9 2002,
Oct 30 2004

BSE
Geophysical
model : 16 TW

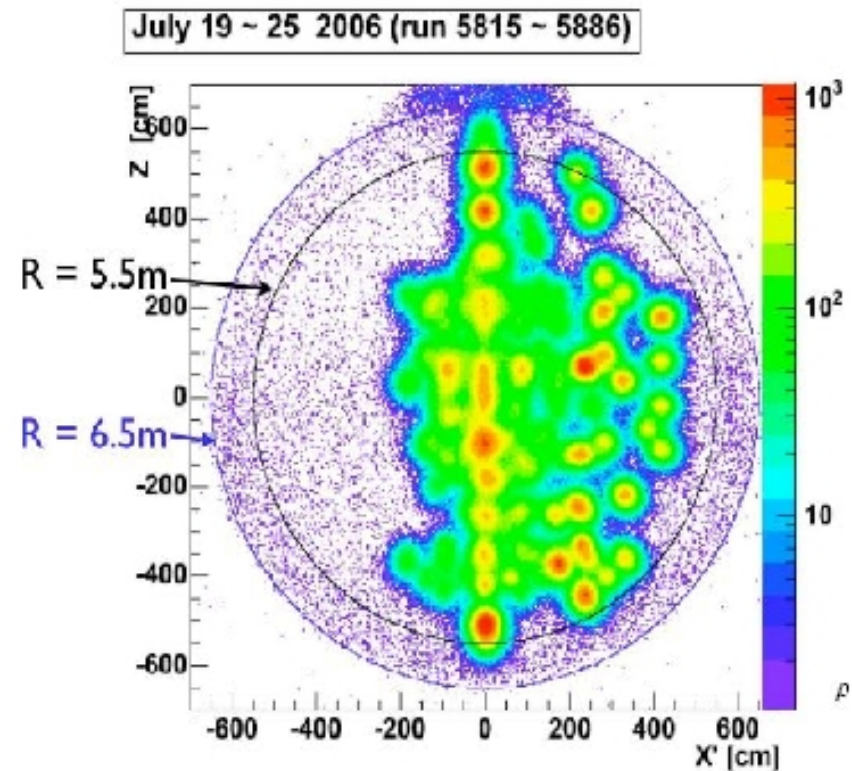
Nature 436, 499 (2005)

Calibration improvements

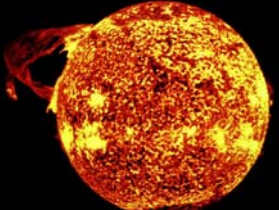
4pi calibration system



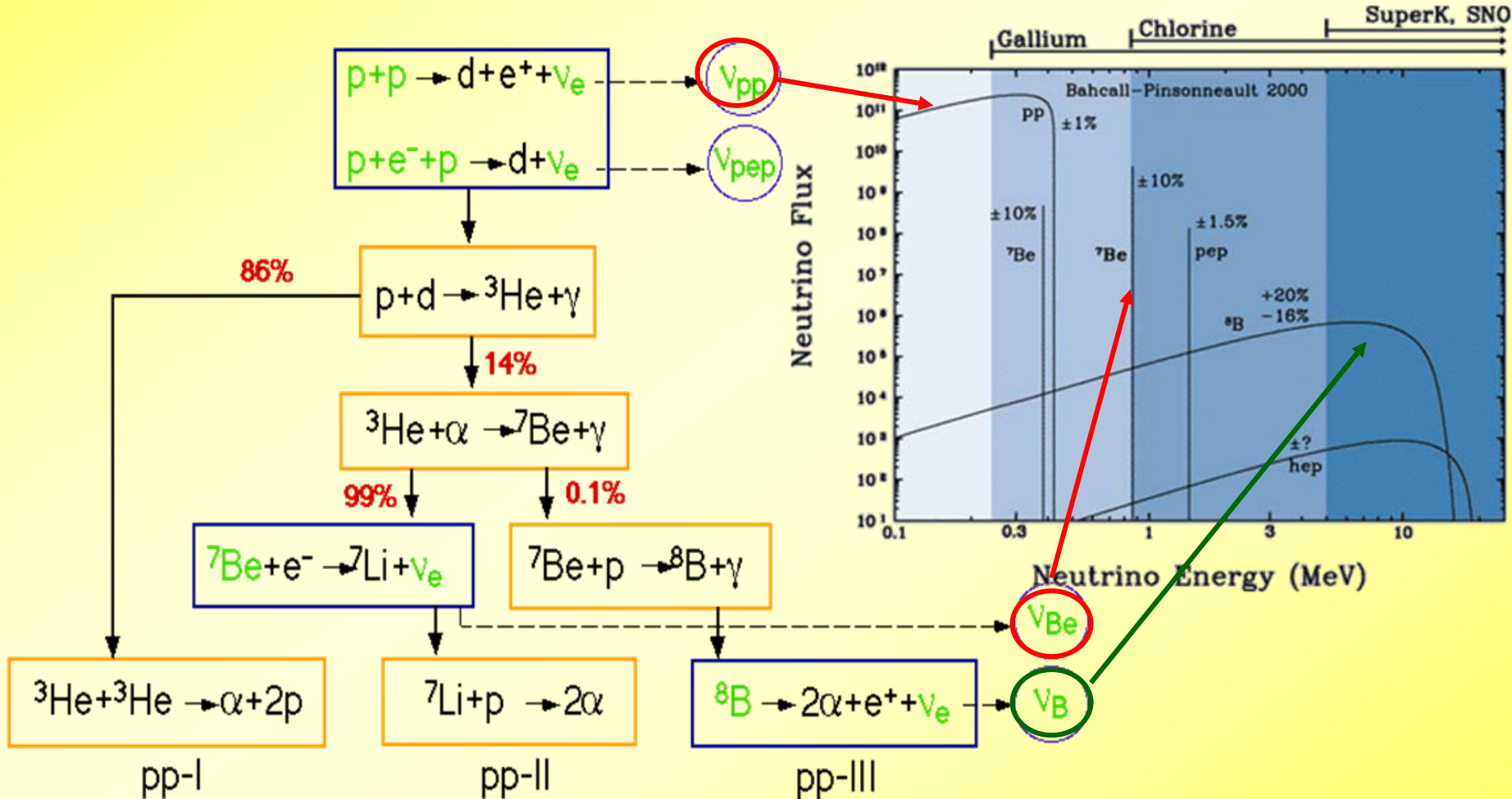
^{60}Co calibration data

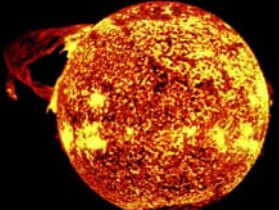


Systematics on anti-neutrino reactor 6.5 % \rightarrow 4%



Solar Neutrinos





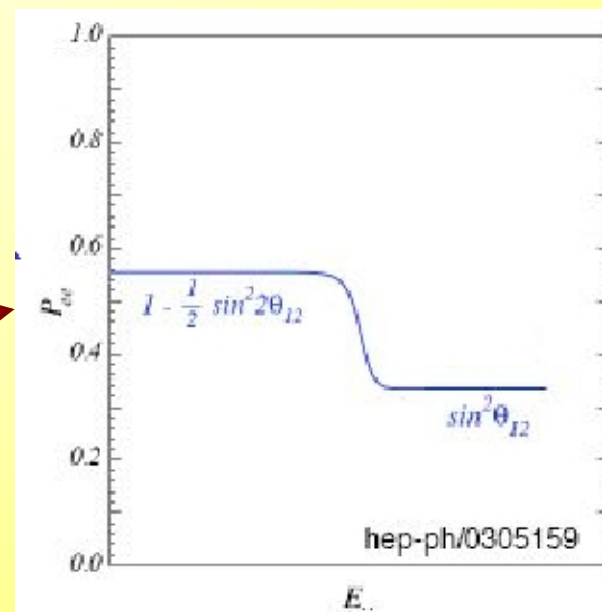
${}^7\text{Be}$ neutrinos

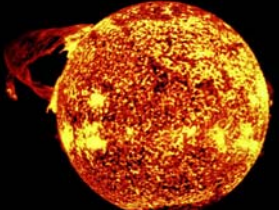
Experimental uncertainty: 40 %

Real-time measurement

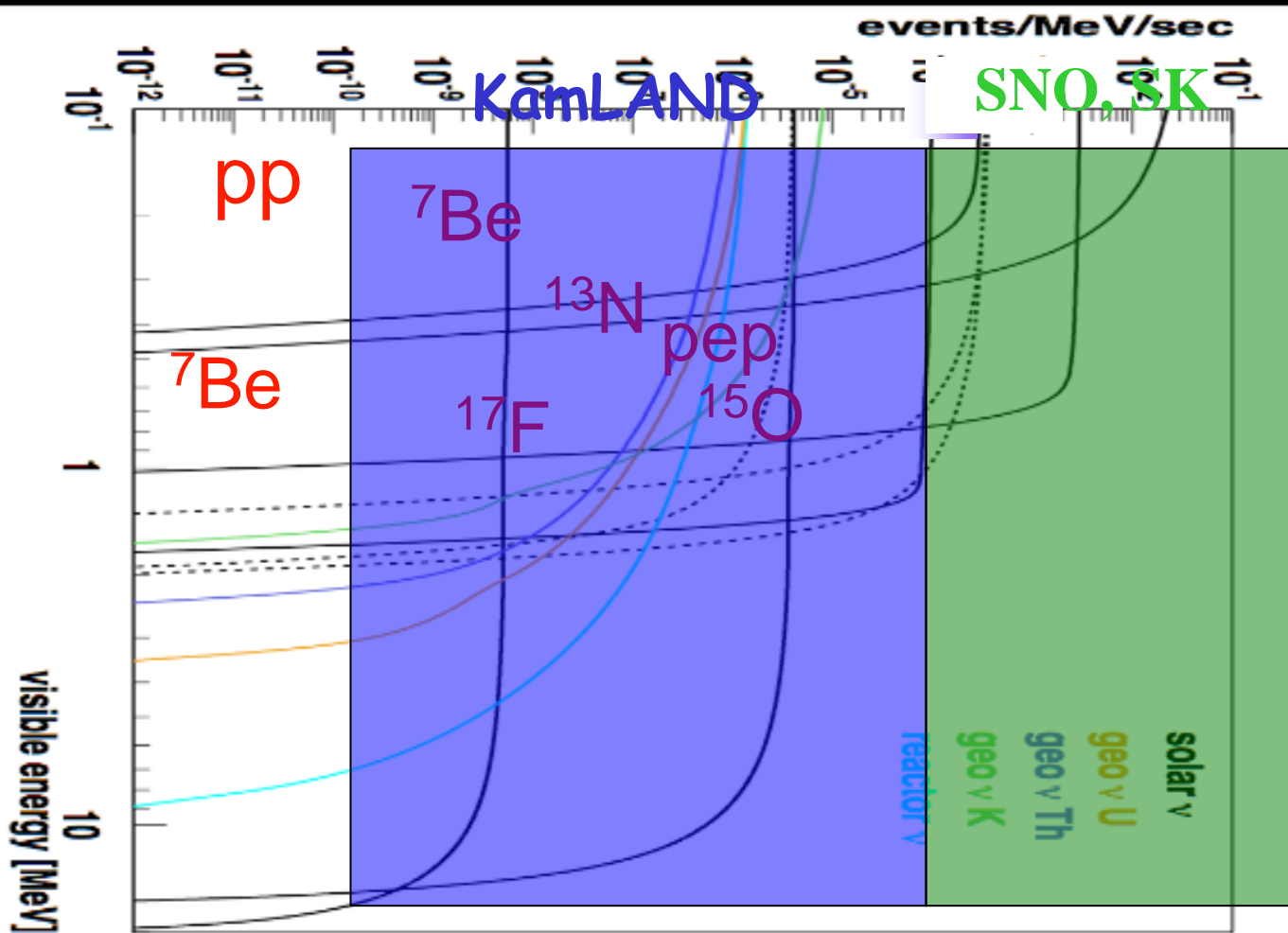
$$\frac{2\Phi({}^7\text{Be})}{(\Phi(\text{pp})-\Phi({}^7\text{Be}))} = \frac{\langle {}^3\text{He}+{}^4\text{He} \rangle}{\langle {}^3\text{He}+{}^3\text{He} \rangle}$$

${}^7\text{Be}$ neutrinos below MSW threshold
Check of MSW effect

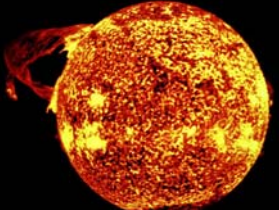




Solar Neutrinos

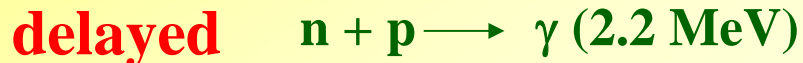
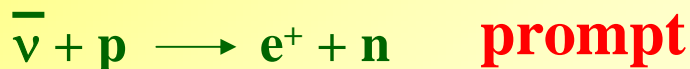


Expected solar yield: $340 \nu / \text{Kt/day}$ (280 keV – 800 keV)

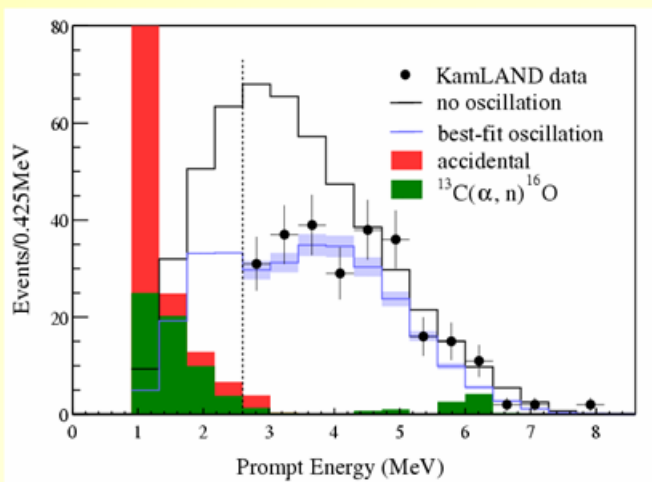


Solar Neutrino detection

Anti-neutrino detection



Powerful signature

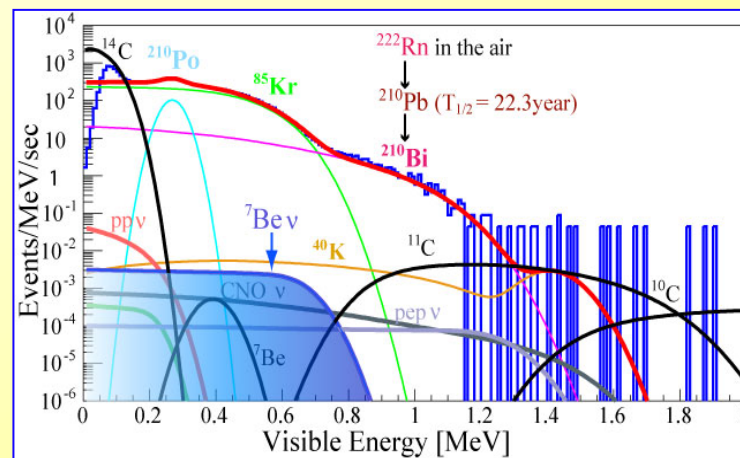


Small background

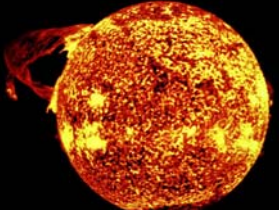
Neutrino detection



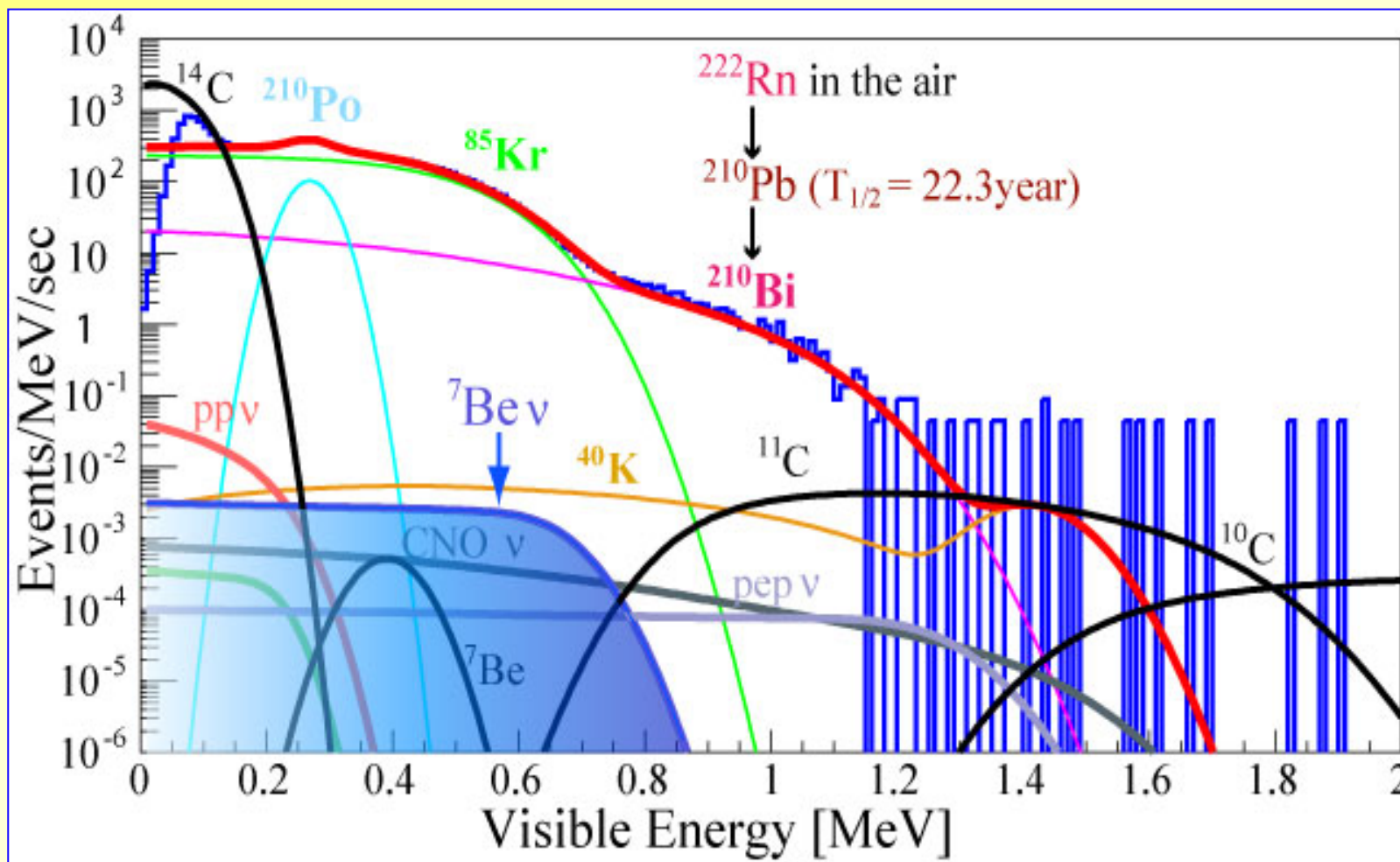
Unspecific Signature

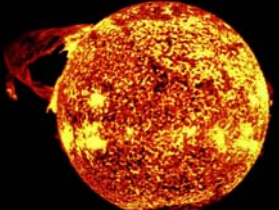


High level of background



Internal Backgrounds



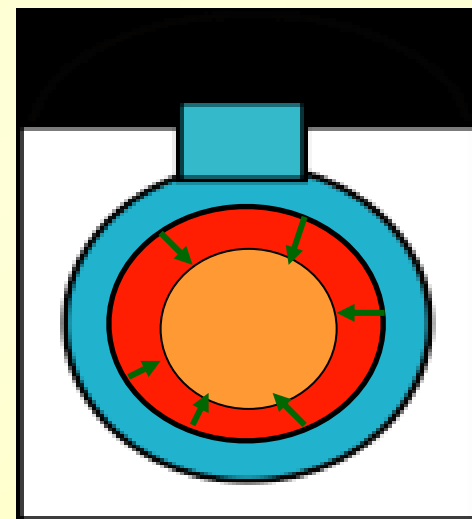


External Backgrounds

External background : ^{40}K and ^{208}Tl

Supressed by fiducial cut $R < 4\text{ m}$

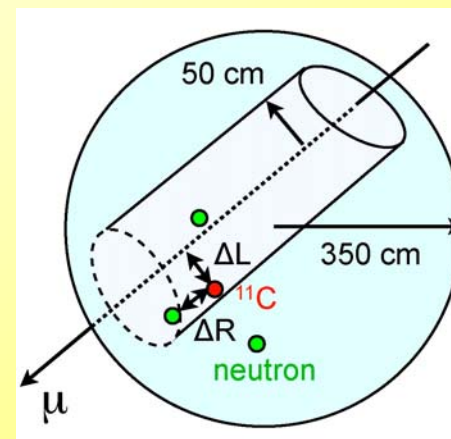
Expected rates 200 x lower than $\nu_{(7\text{Be})}$ signal

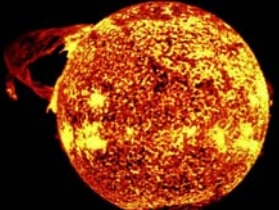


Cosmogenic induced by muons :

^{11}C , ^{12}B , ^7Li , ^7Be ,...

Tagged by veto and in some case
by delayed neutron





Purification goals

Measured Activities in KamLAND

	$T_{1/2}$	Current KamLAND Concentrations	Purification Goal
^{14}C	5730 y	0.5 Bq/m ³	0.5 Bq/m ³ OK
^{210}Pb	22 y	60 mBq/m ³	0.6 $\mu\text{Bq/m}^3$
^{40}K	10^9 y	$1.9 \cdot 10^{-16}$ g/g	10^{-18} g/g
^{85}Kr	11 y	700 mBq/m ³	1 $\mu\text{Bq/m}^3$
^{238}U	10^9 y	$3.5 \cdot 10^{-18}$ g/g	10^{-18} g/g OK
^{232}Th	10^{10} y	$5.2 \cdot 10^{-17}$ g/g	10^{-16} g/g OK



LS Purification

Removal of ^{85}Kr , ^{40}K , ^{210}Pb , ^{210}Bi , ^{210}Po , ^{222}Rn

The KamLAND Collaboration is currently studying the effects of :

- Distillation
- Nitrogen Purging
- Adsorption
- Heating

LS Distillation



Test plant
(Tohoku
University)

^{nat}Kr Reduction: 10^5

Measured by GC

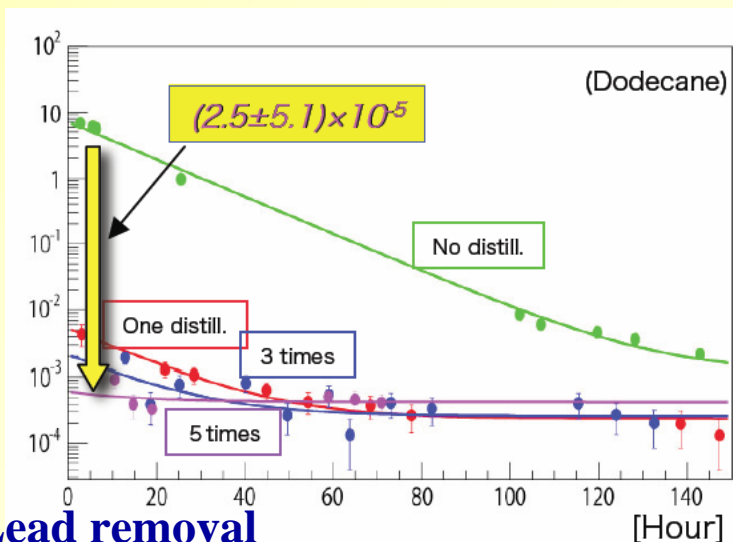
^{222}Rn Reduction: 10^6

Measured by β - α coincidence of
 $^{214}\text{Bi} - ^{214}\text{Po}$ decay ($233 \mu\text{s}$)

^{212}Pb Reduction: 10^4

Measured by β - α coincidence of
 $^{212}\text{Bi} - ^{212}\text{Po}$ decay ($0.43 \mu\text{s}$)

Operates at a 1-2 L/hr



Adsorption



Adsorption is the adherence to a surface.

Adsorption removes charged atoms, i.e. Pb^{+2} by retention on the surfaces of the adsorption particles (silica, Alusil, Cu/Mn).

^{212}Pb Reduction: 30

Measured by Germanium detector or β - α coincidence of $^{212}\text{Bi} - ^{212}\text{Po}$ decay ($0.43 \mu\text{s}$)



Heating



Heating is used to break organo-metallic bonds which then ionize the Pb, Po, Bi, etc atom and can be removed by adsorption or distillation.

Operating Temperature: 100 – 200 °C

Used in combination with distillation or adsorption column. Same removal efficiency seen in both systems

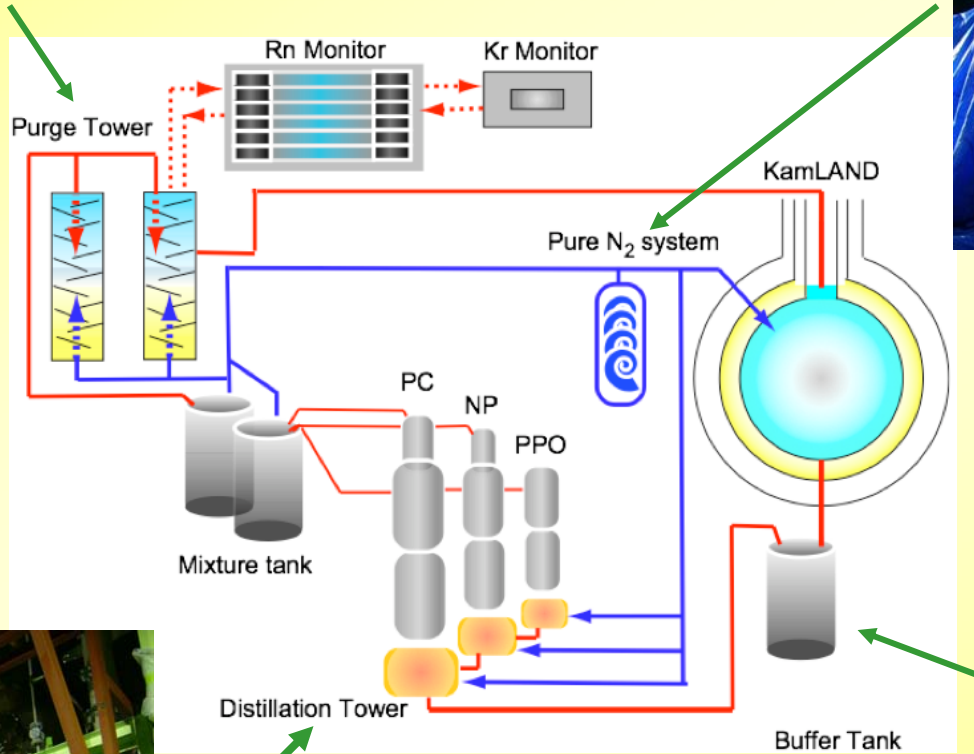
^{212}Pb Reduction boost factor: 10



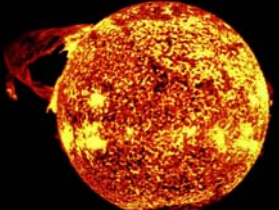
Online monitoring

- Purpose is to insure that we are obtaining high levels of purification and not re-contaminating after purification procedure.
- ^{85}Kr measurement system which will increase our sensitivity to low concentrations by using a cold trap and then passing through an RGA.
- ^{222}Rn measurement (mini-KamLAND)
- Other activities (U, Th, ^{210}Pb) are too low to measure without a detector like KamLAND or long counting times.

Purification line

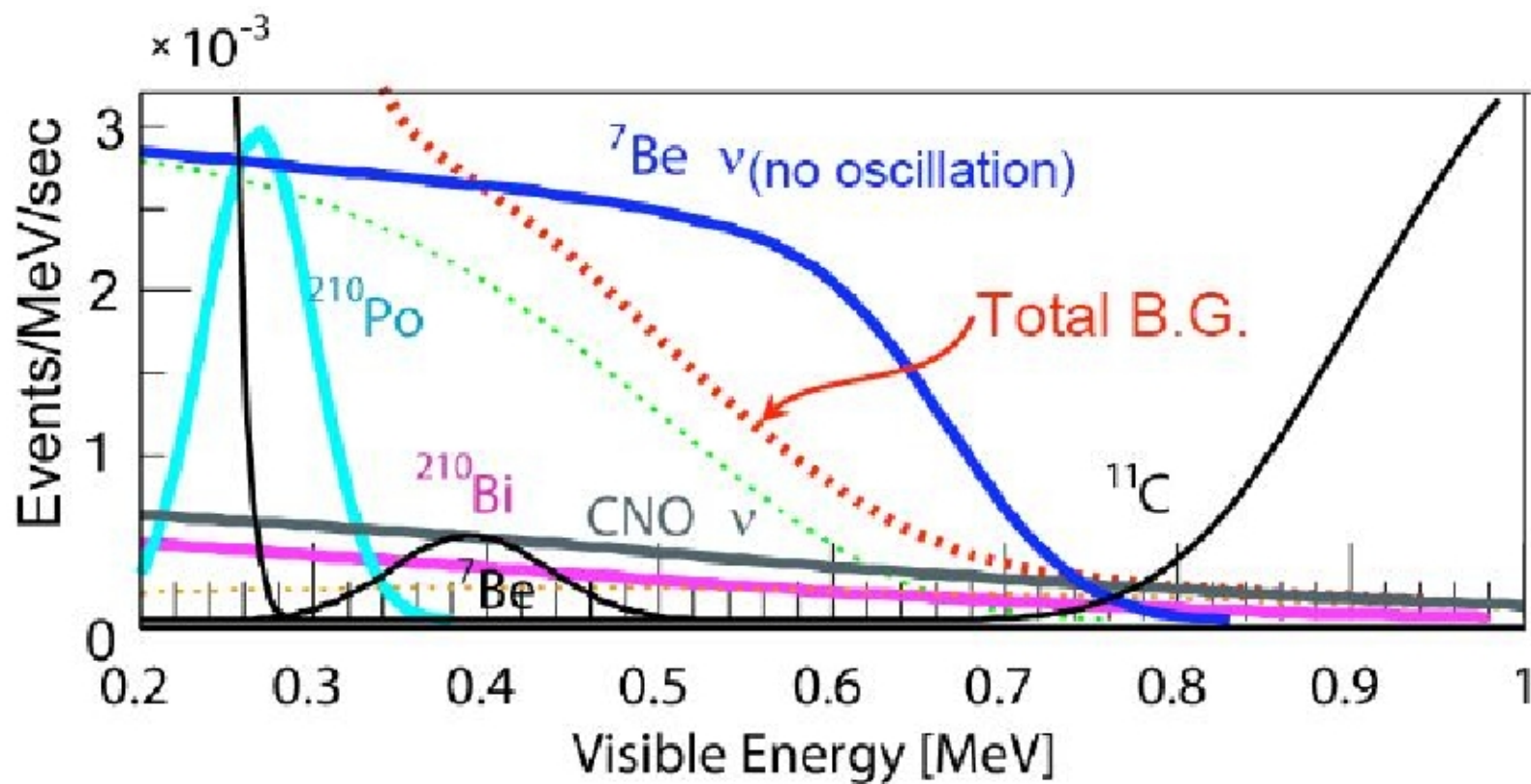


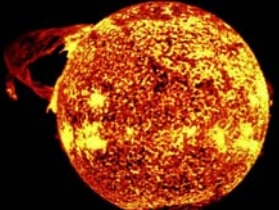
Purification line ready



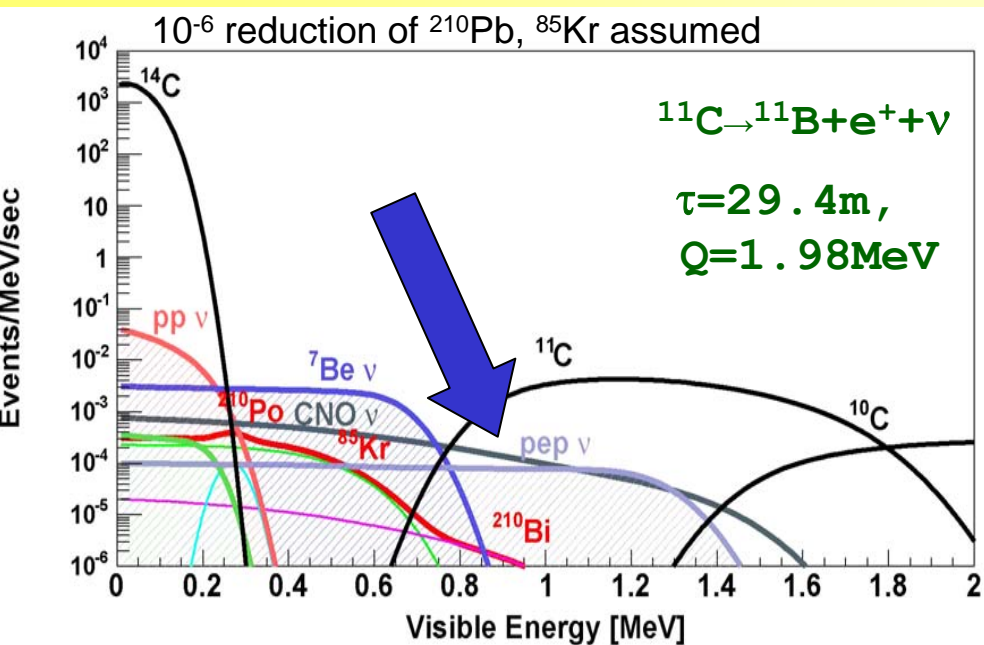
Expected signal for ${}^7\text{Be } \nu$

After $3 \cdot 10^{-5}$ reduction for ${}^{210}\text{Pb}$





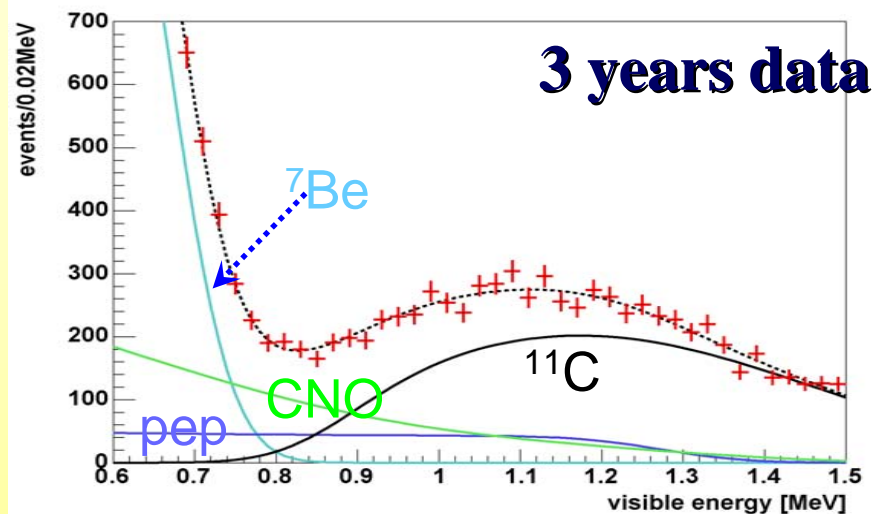
Toward pep/CNO detection

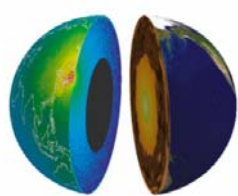


¹¹C reduction by 3-fold coincidence:

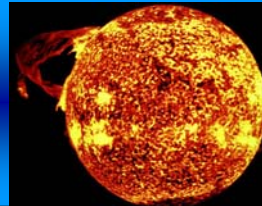
- 1) Muon
- 2) Neutron (2.2MeV
 γ after $\sim 200\mu\text{s}$)
- 3) ¹¹C decay ($\tau = 29.4\text{m}$)

New electronics to detect neutrons
 After large muon signal
 Improvement of muon fitter and
 Muon tracking





Summary



Anti-neutrino reactor:

- Spectral distortion
- Data taking ongoing
- Update of data with full volume calibration soon

Geoneutrinos:

- First detection of geoneutrinos
- Effort to reduce systematic error on background

Solar neutrinos:

- LS purification line is ready
- Goal is to measure 7Be neutrino flux within 10%
- Backgrounds improvements for reactor anti-neutrino and geoneutrinos
- Studies for pep neutrinos detection